

Hybrid (Hamiltonian) Monte Carlo

- We would like to increase the acceptance rate to improve efficiency, and explore the target distribution efficiently ('good mixing')
- We have a hard problem in many dimensions. Solution:
- Make things harder: add in M auxiliary variables, one for each parameter in the model.
- Imagine each of the parameters in the problem as a coordinate.
- Target distribution \rightarrow effective potential
- For each coordinate HMC generates a generalised momentum.
- It then samples from the extended target distribution in $2M$ dimensions.

HMC

- HMC explores this $2M$ -dimensional space by treating the problem as a dynamical system, and evolving the phase space coordinates by solving the dynamical equations.
- Finally, it ignores the momenta (marginalising, as in MCMC), and this gives a sample of the original target distribution.
- May help with decorrelating the points in the chain.
- Invented by particle physicists (Duan et al 1987)

Theory

- Potential $U(\theta) = -\ln p(\theta)$
- For each θ_α , generate a momentum u_α .
- K.E. $K = \mathbf{u}^T \mathbf{u} / 2$
- Define a Hamiltonian

$$H(\boldsymbol{\theta}, \mathbf{u}) \equiv U(\boldsymbol{\theta}) + K(\mathbf{u})$$

- and define an extended target density

$$p(\boldsymbol{\theta}, \mathbf{u}) = \exp[-H(\boldsymbol{\theta}, \mathbf{u})]$$

Magic of HMC

- Evolve as a dynamical system

$$\begin{aligned}\dot{\theta}_\alpha &= u_\alpha \\ \dot{u}_\alpha &= -\frac{\partial H}{\partial \theta_\alpha}\end{aligned}$$



William Rowan Hamilton

- H remains constant, so extended target density is uniform – all points get accepted!
- Also, you can make big jumps – good mixing, if you generate a new \mathbf{u} each time

Complications

- Evolving the system takes time. **Take big steps.**
 - We don't know the complete $U = -\ln p$ (it's what we are looking for)
 - If we can quickly evaluate the derivative of U , fine.
 - We might **approximate U** (from a short MCMC)
- $$U = \frac{1}{2}(\theta - \theta_0)_\alpha C_{\alpha\beta}^{-1}(\theta - \theta_0)_\beta$$
- H is therefore not constant
 - Use Metropolis-Hastings. Accept new point with probability

$$\min \{1, \exp[-H(\boldsymbol{\theta}^*, \mathbf{u}^*) + H(\boldsymbol{\theta}, \mathbf{u})]\}$$

Algorithm

Hamiltonian Monte Carlo

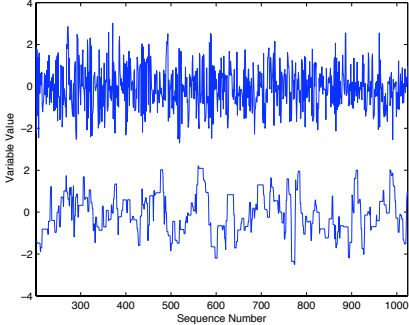
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1: initialize  $\mathbf{x}_{(0)}$ 
2: for  $i = 1$  to  $N_{samples}$ 
3:    $\mathbf{u} \sim \mathcal{N}(0, 1)$ 
4:    $(\mathbf{x}_{(0)}^*, \mathbf{u}_{(0)}^*) = (\mathbf{x}_{(i-1)}, \mathbf{u})$ 
5:   for  $j = 1$  to  $N$ 
6:     make a leapfrog move:  $(\mathbf{x}_{(j-1)}^*, \mathbf{u}_{(j-1)}^*) \rightarrow$ 
 $(\mathbf{x}_{(j)}^*, \mathbf{u}_{(j)}^*)$ 
7:   end for
8:    $(\mathbf{x}^*, \mathbf{u}^*) = (\mathbf{x}_{(N)}^*, \mathbf{u}_{(N)}^*)$ 
9:   draw  $\alpha \sim (0, 1)$ 
10:  if  $\alpha < \min\{1, e^{-(H(\mathbf{x}^*, \mathbf{u}^*) - H(\mathbf{x}, \mathbf{u}))}\}$ 
11:     $\mathbf{x}_{(i)} = \mathbf{x}^*$ 
12:  else
13:     $\mathbf{x}_{(i)} = \mathbf{x}_{(i-1)}$ 
14: end for

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From Hajian 2006

HMC vs MCMC



Typical speed-ups: factor 4.

Supernova distances



- Luminosity Distance depends on Ω_m and H_0 (for flat Universe)

$$f = \frac{L}{4\pi D_L^2}$$

$$D_L = \frac{c(1+z)}{H_0} \int_0^z \frac{dz'}{\sqrt{\Omega_m(1+z')^3 + 1 - \Omega_m}}$$

